

Dynamic Shack-Hartmann System “RAPID”

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UNAM

PAPIIT IN104910

Telescopes OAN

OAN-SPM

- 2.1m telescope
- 1.5m telescope
- 0.84m telescope
- OAN-Tonanzintla
 - 1m telescope

2.1m telescope



Astronomical site SPM

Observing statistics

1982-2002 (Tapia+ 2007)

- “Photometric” nights: 64.1%
- “Spectroscopic” nights: 80.2%
- “Bad weather” nights: 21.2%

• Sky Transparency

Extinction coefficients:

- Visible: $k_V = 0.14$
- Red: $K_R = 0.05$

SPM confirmed have largest fraction of clear nights of any site in the Northern Hemisphere

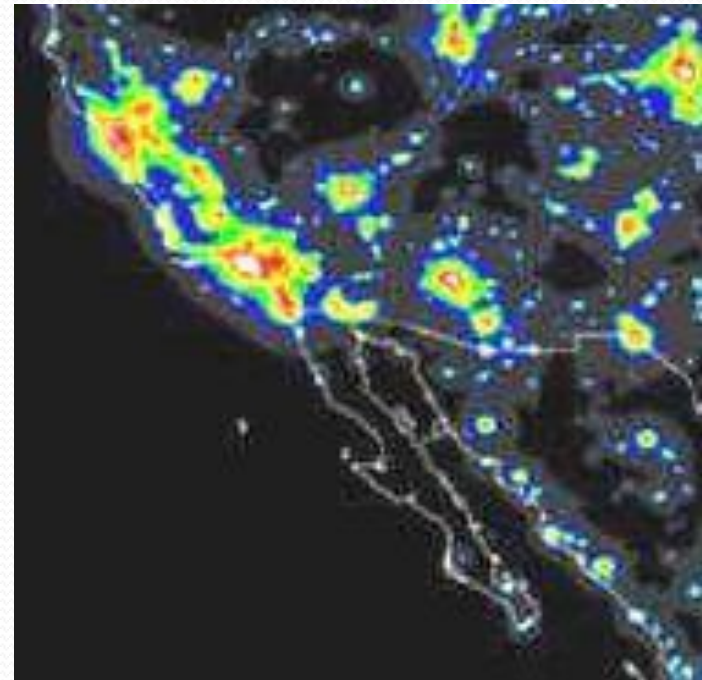
0.84m telescope



Sky Brightness in SPM

Brightness magnitudes

- $m_W = 21.7$
- $m_B = 22.4$
- $m_V = 21.5$
- $m_R = 20.7$
- $m_U = 19.2$



Astronomical site SPM

San Pedro Martir

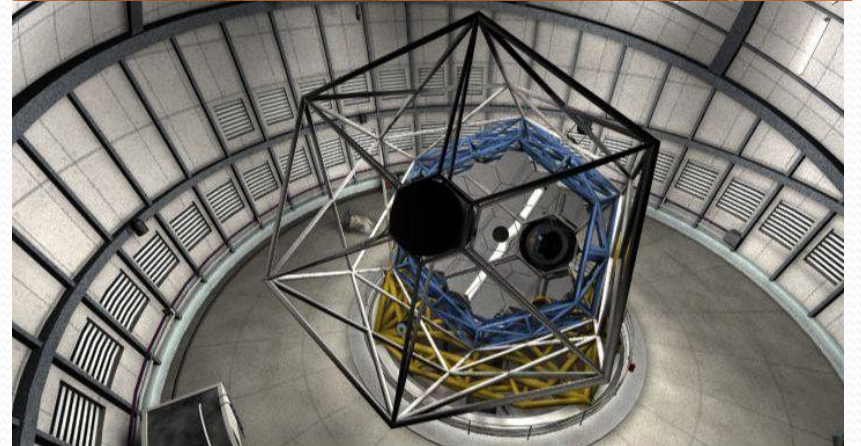
1.5m telescope

- Good sky transparency
- Big number of clear nights
- Excellent image quality
- Photometric stability
- SPM is one of the darkest sites in the Northern Hemisphere
- SPM is good for IR observing
- So, we are looking for big telescopes



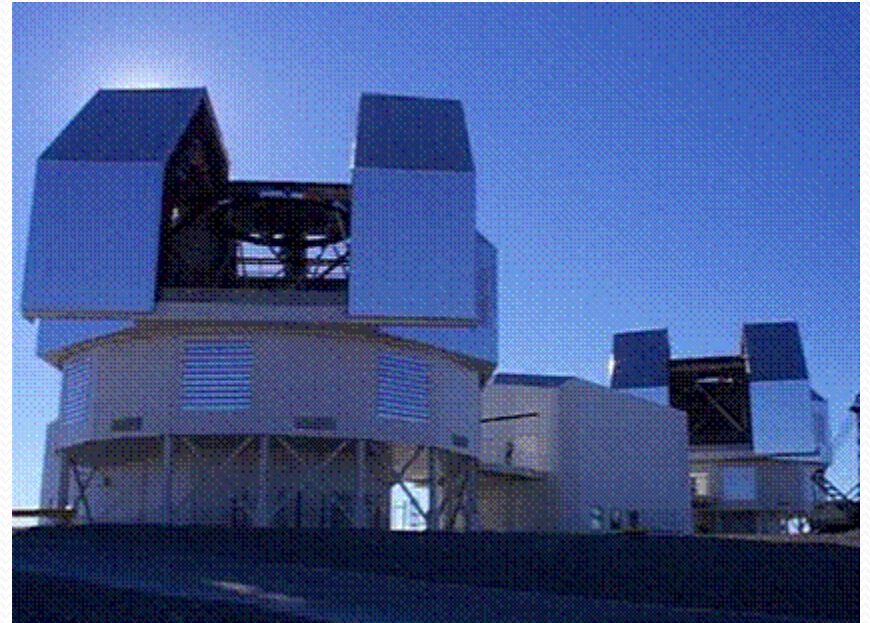
Telescopio Optico-Infrarrojo Mexicano (TIM)

- alt-azimuth mount
- Segmented mirror
- 6.5m > BTA
- (1996-2004)



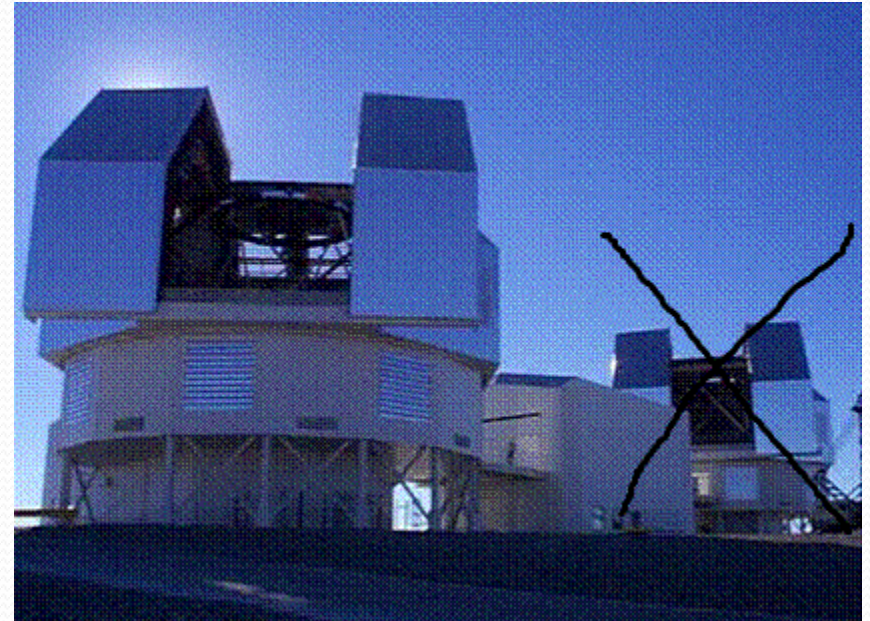
SPM-Twin

- Copy of Magellan
- 2 x 6.5m telescopes f/4.5 :
 - WF-Spectroscopic Telescope
 - WF AO telescope
- *(2005-2008)*



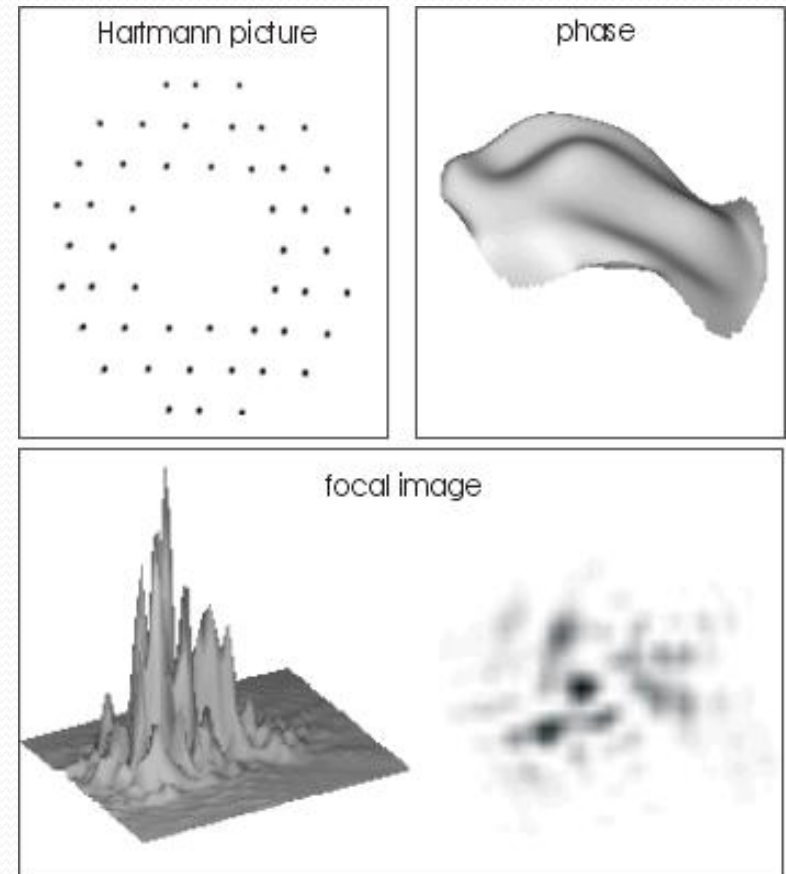
Synoptic All-Sky Infrared Survey (SASIR)

- Magellan-inspired
- 6.5m mirror
- At least we have the mirror.
- (2008-...)



Rapid - Dynamic Shack-Hartmann

- All modern telescopes demand installation of adaptive system.
- So, we have to study of spatial and temporal characteristics of the wavefront distortions.
- We have chosen a Shack-Hartmann wavefront sensor :
 - It is simply
 - Compact data



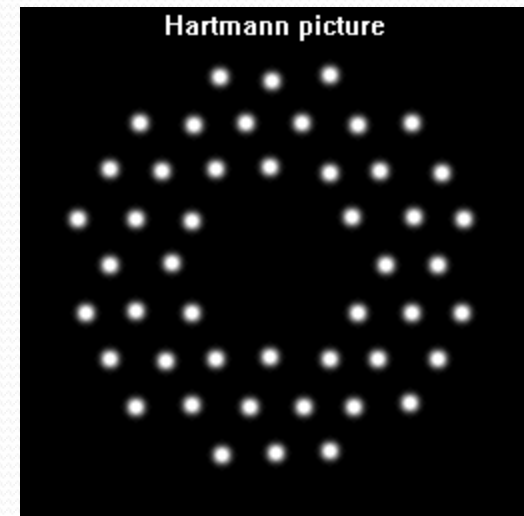
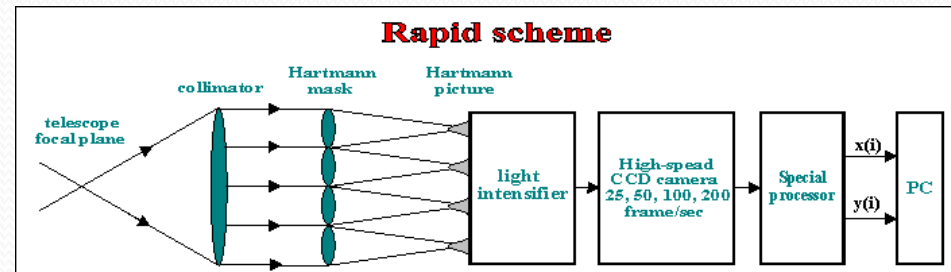
Rapid - Dynamic Shack-Hartmann

1. Collimator:

- Focal distance 200mm

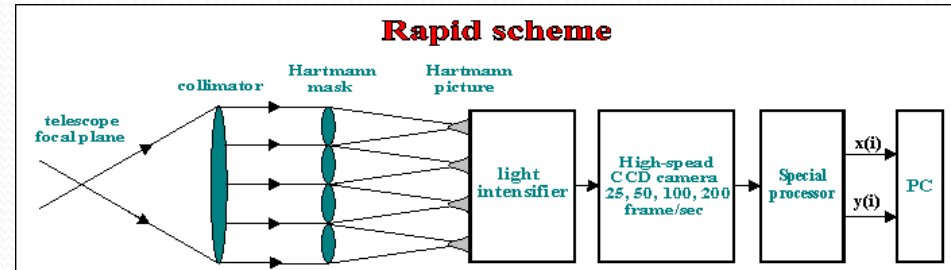
2. Lenslet-Array. Shack-Hartmann diaphragm:

- Consists of 48 lenses with a focal distance of 150 mm
- Hexagonal arrangement to minimize the wavefront measurement errors



Rapid - Dynamic Shack-Hartmann

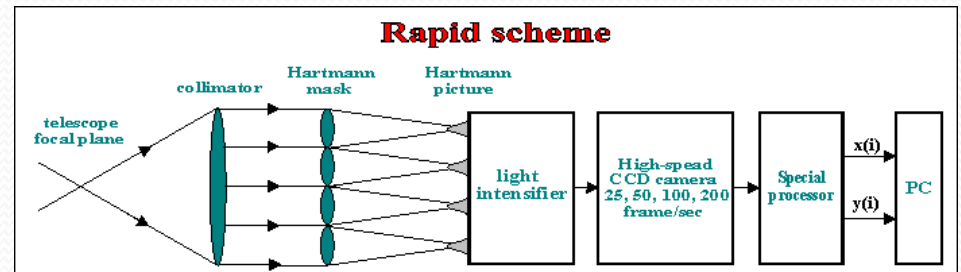
- Image Intensifier Tube
- Generation 2 (ЭП-10)
- Photocathode diameter, mm 24.5
- Resolution, lp/mm 33,8
- Photocathode sensitivity, $\mu\text{A}/\text{lm}$ 280
- Light gain, 30,000
- Dark background brightness, max, cd/m^2 2.5×10^{-3}
- Signal-to-noise ratio 4
- Consumption, max, mA 35
- Mean time before failure, hours 2,000



Rapid - Dynamic Shack-Hartmann

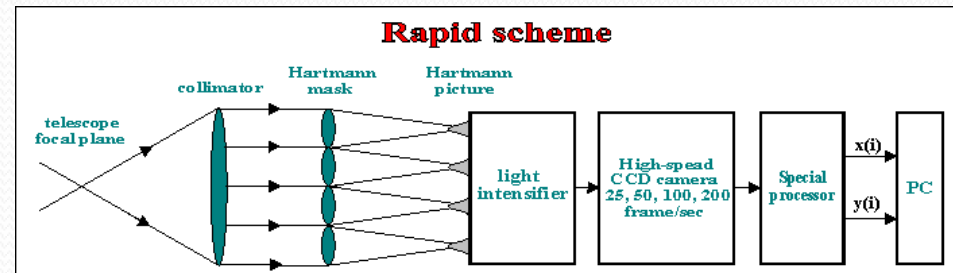
• Light intensifier

- Generation 2 (ЭП-10) 3
- Photocathode operating diameter, mm 24.5 18
- Resolution, lp/mm 33,8 58
- Photocathode sensitivity, $\mu\text{A}/\text{lm}$ 280 1500
- Light gain, 30,000 42,000
- Dark background brightness, max, cd/m^2 2.5×10^{-3} 0.4
- Signal-to-noise ratio 4 21
- Consumption, max, mA 35 20
- Mean time before failure, hours 2,000 10,000



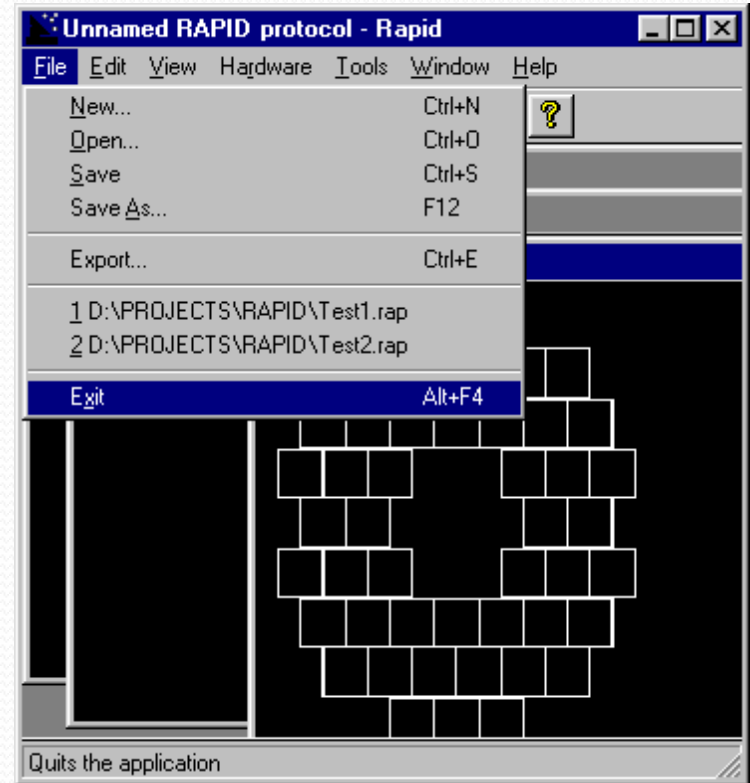
Rapid - Dynamic Shack-Hartmann

- High-speed CCD
 - Based on SONY ICX055AL 1/3-inch CCD Image Sensor for CCIR B/W Camera
 - 25,50,100,200 fr./sec.
 - 200 is possible only for 256*288
 - Electronic shutter
- Special processor
 - Calculate coordinates of the center of mass for each spot.



Rapid - Dynamic Shack-Hartmann

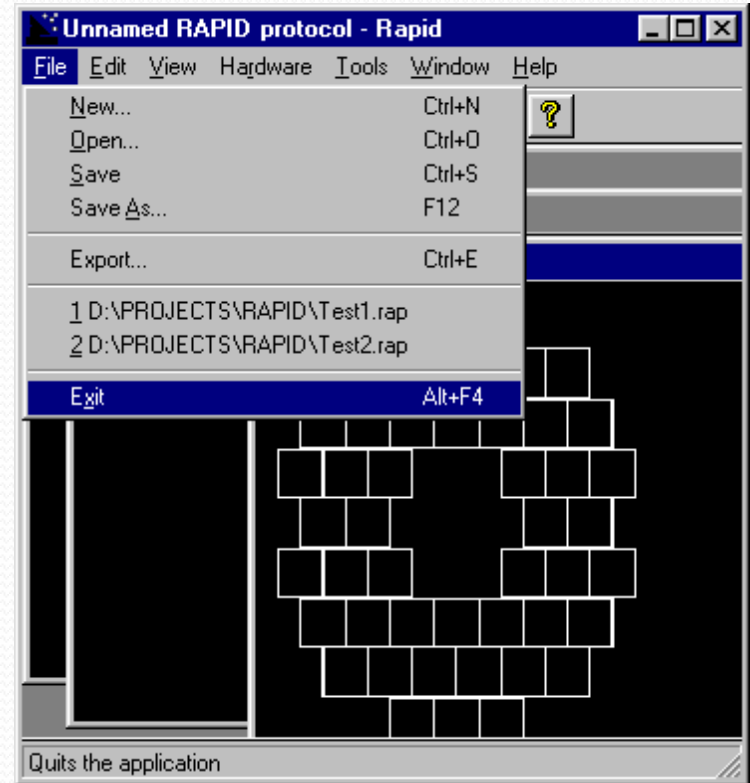
- Rapid Software:
 - Creation, saving or loading of RAPID protocols
 - Viewing and addition copies of the camera image, with tuning and textual comments saving
 - Execution of operation of data accumulation in "Master" (Wizard) mode .



steps

First step:

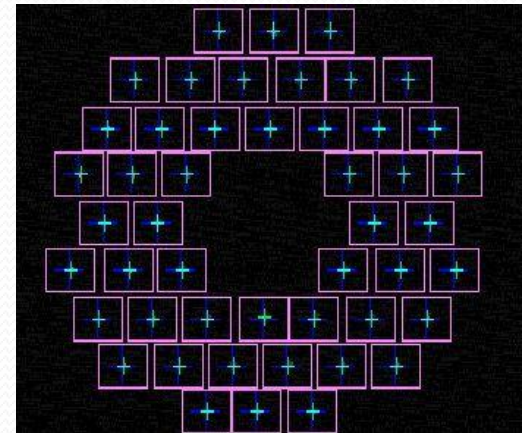
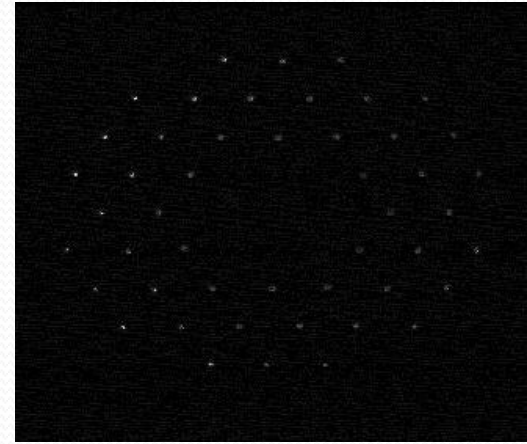
- choose of parameters of the camera
- assignment of the dark current.
- choose a video mode for reading data
- amount of frames, used for a dark level calculation.
- Shutter on/off
- amplification



steps

Second step:

- Parameters for Hartmann picture place calculation : frequencies
 - 200/100/50 (for 256*288)
 - 100/50/25 (for 512*288)Hertz are available,
- amount of averaged frames: from 2 up to 1024
- Amplification of the camera
- Amplification of the II
- "Preview" for previewing the image.



steps

Third step:

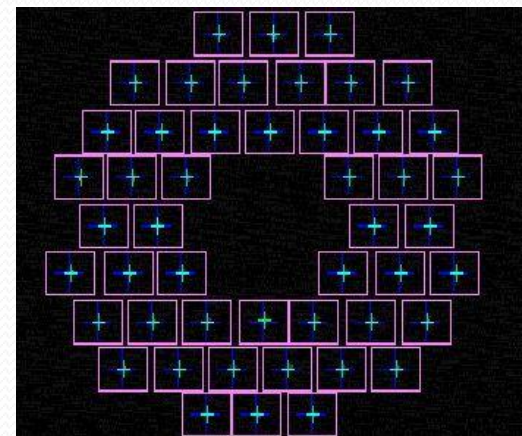
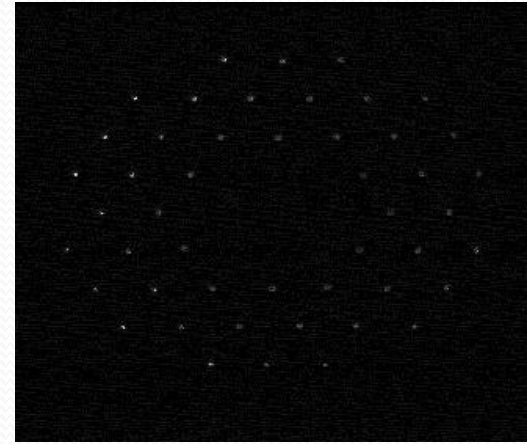
Choice of threshold meanings
for Hartmann picture
calculation. appropriate
buttons

Fourth step:

Save data on the memory

Fifth step:

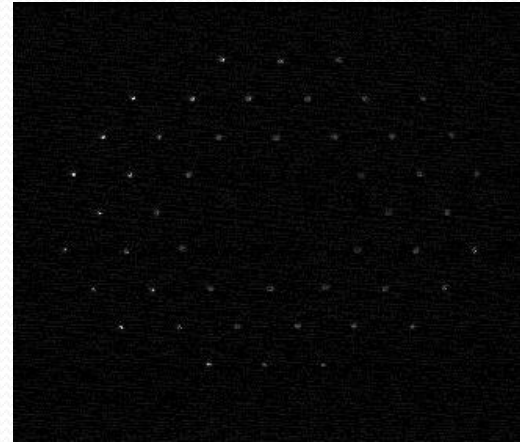
Save on the hard disk

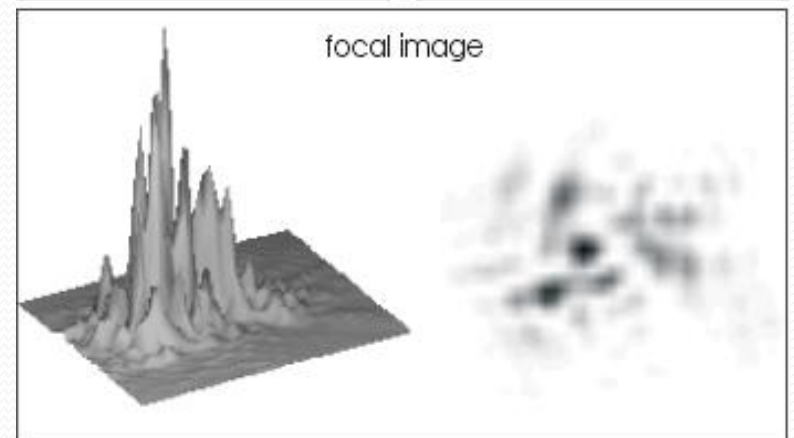
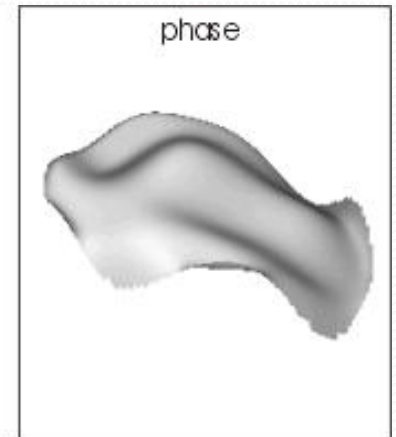
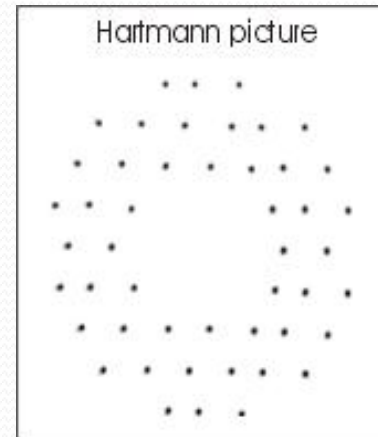


Rapid Software

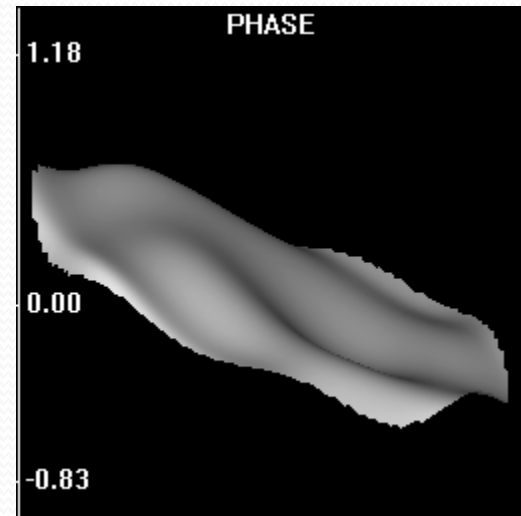
In the same
protocol also it is
possible:

- write text of the comment
- save up to 50 mages
directly from ccd.

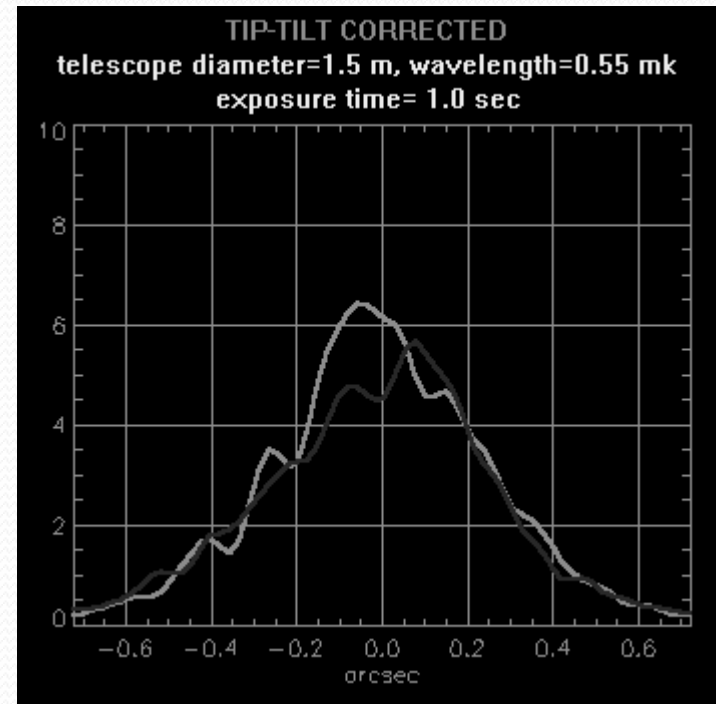




- The phase reconstruction is very fast and simple.



Hartmann method



TIP-TILT, DEFOCUS, ASTIGMATISM CORRECTED
telescope diameter=1.5 m, wavelength=0.55 μ m
exposure time= 1.0 sec

